

## Imaging a single atom in a time-of-flight experiment

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# **New Journal of Physics**

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## Imaging a single atom in a time-of-flight experiment

## A Fuhrmanek<sup>1</sup>, A M Lance<sup>2</sup>, C Tuchendler<sup>1</sup>, P Grangier<sup>1</sup>, Y R P Sortais<sup>1,3</sup> and A Browaeys<sup>1</sup>

 <sup>1</sup> Laboratoire Charles Fabry, Institut d'Optique, CNRS, Université de Paris-Sud, Campus Polytechnique, 2 avenue Augustin Fresnel, RD 128, 91127 Palaiseau Cedex, France
<sup>2</sup> Joint Quantum Institute, National Institute of Standards and Technology and University of Maryland, Gaithersburg, MD 20899, USA
E-mail: yvan.sortais@institutoptique.fr

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**Abstract.** We perform fluorescence imaging of a single <sup>87</sup>Rb atom after its release from an optical dipole trap. The time-of-flight expansion of the atomic spatial density distribution is observed by accumulating many single atom images. The position of the atom is revealed with a spatial resolution close to 1  $\mu$ m by a single-photon event, induced by a short resonant probe. The expansion yields a measure of the temperature of a single atom, which is in very good agreement with the value obtained by an independent measurement based on a release-and-recapture method. The analysis presented in this paper provides a way of calibrating an imaging system useful for experimental studies involving a few atoms confined in a dipole trap.

<sup>3</sup> Author to whom any correspondence should be addressed.

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#### 1. Introduction

Time-of-flight imaging of ultra-cold atomic gases in expansion is a common way to study their properties. It provides a direct measurement of the momentum distribution and is therefore routinely used to extract the temperature of cold thermal samples [1]. It can also give access to spatial density or momentum correlations in atomic ensembles. These features have, for instance, enabled the observation of bunching (anti-bunching) with bosonic (fermionic) atoms [2, 3]. They also enable the study of condensed matter phenomena that emerge when confining matter waves in periodic optical potentials [4, 5]. Ultimately, one would like to observe the atoms individually in these mesoscopic systems, not only when the atoms are confined but also when they move or are released from the trap in order to access out-ofequilibrium properties. Although fluorescence imaging is widely used in experiments to detect single trapped atoms [6]–[8], and sometimes spatially resolve them [9]–[11], fluorescence imaging of freely propagating single atoms has been demonstrated only recently [12]. In that experiment, cold atoms are released from a trap and fall under gravity through a sheet of light, which is imaged on an intensified CCD camera using efficient collection optics. The presence of an atom is revealed by an individual spot corresponding to the detection of many fluorescenceinduced photons. The detection efficiency of a single atom is close to unity and the spatial resolution ( $\leq 10 \,\mu$ m) is set by the motion of the atom in the light sheet.

In this paper, we present a complementary approach where we demonstrate time-of-flight fluorescence imaging of a single Rb atom in free space, based on single-photon detection, with a spatial resolution of  $\sim 1 \,\mu$ m. A single atom is first trapped in a microscopic dipole trap and then released in free space where it evolves with its initial velocity. To detect the atom and locate it with the best accuracy possible, we illuminate it with a very short pulse of resonant light and collect the fluorescence on an image intensifier followed by a CCD camera. The presence of the atom is revealed by a single-photon event and, for probe pulses as short as 2  $\mu$ s, the probability to detect an atom in a single shot of probe light is 4.4%. We repeat the experiment until the spatial distribution of the atom is reconstructed with a sufficient signal-to-noise ratio; the accumulation of successive single-atom images yields an average result that exhibits the same features as would a single experiment with many non-interacting atoms. Average images recorded for increasing times-of-flight allow us to measure the root mean square (abbreviated rms) velocity

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of the atomic expansion, and thus the temperature of a single atom. Our method allows us to measure temperatures over a wide range.

The analysis of the time-of-flight of a single atom released from our microscopic dipole trap also serves as a calibration of our imaging system. This calibration will be used in future experiments where we plan to study the behavior of a cloud of a few tens of cold atoms held in the microscopic trap. In this regime, the cloud is very dense and light scattering of near-resonant light used for diagnostic purposes may exhibit collective behavior (see e.g. [13]). It is therefore important to understand the optical response of the imaging system in the single-atom case to interpret the images in the multi-atom regime where collective effects may come into play. Moreover, because the atoms can be illuminated right after their release from the dipole trap, our method allows us to explore the properties of the momentum distribution of such a gas in the near-field regime.

The paper is organized as follows. In section 2 we describe the experimental setup. In section 3 we explain the requirements to perform single-atom time-of-flight imaging. Section 4 describes the experimental sequence and shows images of a single atom taken by a CCD camera after variable time-of-flight. We extract the temperature using a standard fit based on an expansion model. Section 5 details the effects that contribute to the spatial resolution of our imaging system. Section 6 re-analyzes the data by using a Monte Carlo simulation taking into account the effects described in section 5. The temperature result is also compared to an independent measurement based on a release-and-recapture technique. Finally, section 7 examines the noise sources in our imaging system.

#### 2. Experimental setup

Our experimental setup has been described elsewhere [9] and is summarized in figure 1. Briefly, a single <sup>87</sup>Rb atom is trapped in a tight optical dipole trap. The dipole trap is produced by focusing a laser beam ( $\lambda_{trap} = 850 \text{ nm}$ ) down to a spot with waist  $w_0 = 1.1 \mu \text{m}$ , using a high numerical aperture aspheric lens (NA = 0.5). The trap depth can be as large as 20 mK for a laser power of 80 mW. We use the same lens to collect the fluorescence light ( $\lambda_{fluo} = 780 \text{ nm}$ ), which is sent onto an avalanche photodiode (APD) and an image intensifier<sup>4</sup> followed by a low-noise CCD camera<sup>5</sup> (see below for more details).

In order to image the atom, we illuminate it with probe light, which consists of two counterpropagating beams (to avoid radiation pressure force) in a  $\sigma_+-\sigma_-$  configuration, and is resonant with the (5<sup>2</sup>S<sub>1/2</sub>, F = 2) to (5<sup>2</sup>P<sub>3/2</sub>, F' = 3) transition. The saturation parameter of the probe light is  $s = I/I_{sat} \sim 1$  for each beam. We also superimpose repumping light on the probe beams, tuned to the (5<sup>2</sup>S<sub>1/2</sub>, F = 1) to (5<sup>2</sup>P<sub>3/2</sub>, F' = 2) transition.

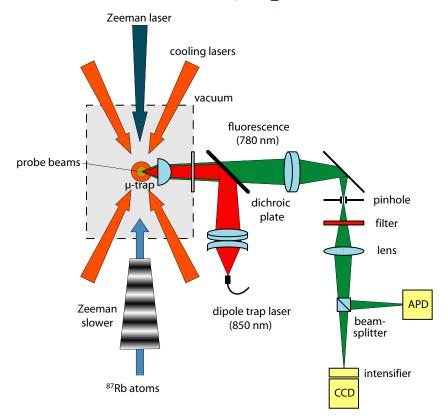
#### 3. Requirements for time-of-flight imaging of a single atom

The principle of a time-of-flight experiment is to measure the position of atoms after a period of free expansion. From the rms positions of the atoms, one extracts the rms velocity  $\sigma_v$  of the atoms. In our experiment, we measure the position of the atom by illuminating it with

<sup>4</sup> Model C9016-22MGAAS from Hamamatsu. The phosphor screen of the intensifier is then imaged onto the CCD camera through a 1 : 1 relay lens.

<sup>5</sup> Model Pixis 1024 from Princeton Instruments.

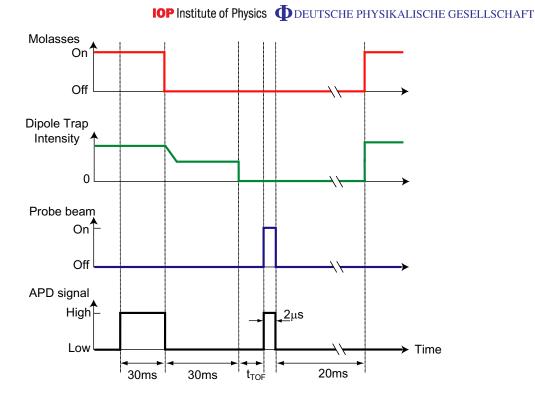
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**Figure 1.** Optical setup of our trapping (red) and imaging (green) systems. A single atom is trapped at the focal point of the aspheric lens and emits fluorescence photons when illuminated with probe beams (perpendicular to the plane of the figure). A fraction of the emitted fluorescence is sent to a fiber-coupled APD in a photon counting mode, which is used to trigger the time-of-flight sequence. The remaining fluorescence is also detected by an intensified CCD camera (I-CCD), with an efficiency  $\eta_d = 2.2 \times 10^{-3}$ , taking into account the solid angle of the aspheric lens (6.7%), the transmission of the optics (33%) and the measured quantum efficiency of the intensifier photocathode ( $\eta_{\text{intensifier}} = 10\%$ ).

resonant laser light and collecting its fluorescence. This method requires that the position of the atom change by less than the resolution of the imaging system during the light pulse. In our case, the imaging system is diffraction limited with a resolution  $\sigma_{\text{diff}} = 0.5 \,\mu\text{m}$ . This imposes a pulse duration of  $\tau < \sigma_{\text{diff}}/\sigma_v$ . Typically, for a rubidium atom at the Doppler temperature  $(T_{\text{Doppler}} \simeq 150 \,\mu\text{K})$ , this yields probe pulses as short as  $4 \,\mu\text{s}$ . For a collection efficiency of  $\sim 1\%$  and a scattering rate  $R \approx \Gamma/2 \simeq 2 \times 10^7 \,\text{s}^{-1}$  ( $\Gamma$  is the line width of the optical transition), the number of detected photons per pixel would approach unity in a single shot, which is well below the capabilities of our CCD camera.

We solved this issue by inserting a light intensifier in front of the CCD camera (see figure 1). The intensifier acts as a fast shutter (opened during the probe pulse only), and amplifies a single-photon event to a level about two orders of magnitude above the noise level of the CCD camera. Using this intensifier, the presence of one atom is revealed by one single-photon event (the case of detecting more than one photon emitted by a single atom during the probe pulse is very unlikely).

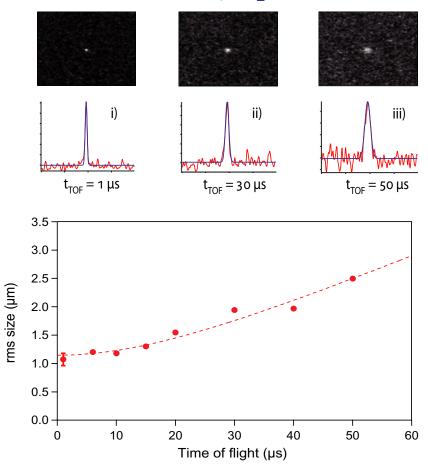


**Figure 2.** Time sequence for a time-of-flight experiment. Note that the time axis is not drawn to scale. The loading sequence lasts  $\sim 1$  s, whereas the adiabatic cooling, time-of-flight and probing sequences are much shorter.

#### 4. Experimental sequence and results

The experimental sequence is summarized in figure 2. It starts with loading and cooling a single atom in the dipole trap. Under illumination by the cooling beams, photons scattered by the trapped single atom are partially collected by the same aspheric lens and directed toward the APD. This light is used as a trigger signal for the subsequent time-of-flight sequence. The cooling beams are switched off immediately upon detection of the atom. The single atom is kept in the dipole trap for an extra 30 ms where the trapped atom can be further cooled by adiabatically ramping down the trap depth [14]. We also use this 30 ms interval to let the atoms in the molasses spread out, with all cooling beams having been switched off. This precaution is taken in order to minimize light scattered by the background molasses during the subsequent probe pulse.

After the single atom is trapped and cooled, the dipole trap is switched off and the singleatom time-of-flight experiment takes place. We let the single atom fly for a variable time  $t_{\text{TOF}}$  and then illuminate it by a 2  $\mu$ s pulse of probe light. At the same time, the intensifier is switched on for 2  $\mu$ s and the probe-induced fluorescence is collected by the intensified CCD camera. The loading sequence is then started again, in order to prepare for the next timeof-flight experiment. The acquisition of one image for a given time-of-flight is performed by repeating the experimental sequence described above, with a cycle rate of ~0.5–2 s<sup>-1</sup>, and accumulating the total fluorescence light on the CCD. When a sufficient number of photons has been detected (typically 100), the CCD chip is read out and the image is displayed. Note that, for each sequence, the CCD receives light only during the 2  $\mu$ s the intensifier is on. In this



**Figure 3.** Results of a typical single-atom time-of-flight experiment. The atom is released from a trap with depth  $U \approx 0.08$  mK. The rms size of the 'cloud' is plotted versus the time-of-flight of the single atom after it is released from the dipole trap. The dashed line is a fit to the data (circles), using equation (1). We show a typical error bar obtained by repeating the same experiment several times. Insets show images and associated cross-sections of the data for three particular times-of-flight. Each image results from the detection of a large number of successively trapped single atoms. The rms size of the cloud is thus the rms position of a single atom after a given time-of-flight. Image (i) corresponds to 3400 sequences and 150 detected photons, and therefore 150 detected atoms. Image (iii) corresponds to 12 000 accumulations and 520 detected atoms.

way the intensifier also serves as a fast switch, preventing stray light from reaching the CCD during the cooling and trapping phases.

Figure 3 shows typical images taken for times-of-flight as long as 50  $\mu$ s. The longer the time-of-flight, the lower the peak signal, and the larger the number of accumulations required. For a measured rms size  $\sigma \simeq 1 \,\mu$ m (corresponding to the time-of-flight  $t_{\text{TOF}} = 1 \,\mu$ s of image (i)), we perform ~3400 sequences, corresponding to 3400 single trapped atoms, and detect 150 photons (this number of photons is extracted from an independent calibration of the intensifier response to a single-photon event). This means that the probability to detect a single atom in a single realization of the experiment is  $4.4 \times 10^{-2}$  when using a 2  $\mu$ s probe.

**Table 1.** Spatial resolution budget of our system, when the light source is a single atom with temperature  $T = 20 \,\mu\text{K}$  illuminated by a probe pulse with duration  $\tau = 2 \,\mu\text{s}$  (see text). The rms size of the global response is the quadratic sum of the different rms contributions.

Effect	rms size (µm)
Intensifier	0.9(2)
Diffraction ( $\sigma_{\text{diff}}$ )	0.5(1)
Atomic thermal distribution ( $\sigma_{\perp}(0)$ )	0.3
Depth of focus ( $\sigma_{\parallel}(0)$ )	0.1
Atomic displacement during $\tau$ ( $\sigma_{\tau,\text{thermal}}$ )	0.04
Atomic random walk ( $\sigma_{\tau, \text{scatter}}$ )	0.02
Quadratic sum	1.1(2)

The images are well fitted by a 2D Gaussian model. Within the error bars the images are isotropic. We plot the rms size  $\sigma$  of the expanding 'cloud' along one axis versus the time-of-flight  $t_{\text{TOF}}$ . We fit the data shown in figure 3 by the general form

$$\sigma(t_{\rm TOF}) = \sqrt{\sigma(0)^2 + \sigma_v^2 t_{\rm TOF}^2}$$
(1)

that gives the rms position of a particle after a time-of-flight  $t_{\text{TOF}}$  when the initial position and the velocity are taken from distributions with standard deviations  $\sigma(0)$  and  $\sigma_v$ . We find  $\sigma(0) = 1.1(1) \,\mu\text{m}$  and  $\sigma_v = 45(2) \,\text{mm s}^{-1}$ . The energy distribution of a single atom in the trap being a thermal Maxwell Boltzmann distribution [14], this translates into a temperature  $T = m\sigma_v^2/k_{\rm B} = 21(2) \,\mu\text{K}$  (*m* is the mass of a rubidium atom and  $k_{\rm B}$  is the Boltzmann constant).

Let us now compare the result for  $\sigma(0)$  to the expected rms radial position of an atom in equilibrium and trapped in a harmonic potential with depth U and transverse size  $w_0$  (at  $1/e^2$ ), i.e.

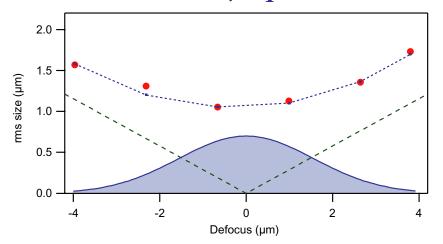
$$\sigma_{\perp}(0) = \sqrt{\frac{k_{\rm B}T}{m\omega_{\perp}^2}},\tag{2}$$

where  $\omega_{\perp}$  is the radial oscillation frequency of the atom in the trap. With  $\omega_{\perp} \approx 2\pi \times 26$  kHz and  $T = 20 \,\mu$ K, we find  $\sigma_{\perp}(0) = 0.3 \,\mu$ m, below the diffraction limit of the imaging system. Taking the latter into account, we should thus expect an rms size of 0.6  $\mu$ m at null time-of-flight, i.e. a factor 1.8 below the actual data.

#### 5. Spatial resolution of our imaging system

In order to understand the size at  $t_{\text{TOF}} = 0$ , we investigated experimentally the effects that contribute to the loss in resolution of our imaging system and lead to the measured value  $\sigma(0)$ . These effects are listed in table 1 and sum up quadratically to yield a value of 1.1  $\mu$ m, in agreement with the measure of  $\sigma(0)$  obtained in section 4.

The dominant contribution comes from the loss of resolution of the imaging system because of the intensifier. This contribution was measured by imaging an object with a sharp edge and characterizing the blurred edge in the image obtained. The second largest contribution



**Figure 4.** rms size of the image of a single trapped atom for various positions of the trap along the optical axis of the imaging system. For small values of the defocus, the image size is given essentially by the response of the intensifier. For large values of the defocus, the main contribution comes from the depth of focus, which scales linearly with defocus according to equation (3) (dashed line). Our data (red circles), which agree well with a Monte Carlo simulation (blue dotted line), result from the convolution of both the optical response of the imaging system and the longitudinal profile of the thermal distribution (blue filled curve) with a width  $\sigma_{\parallel}(0) = 1.6 \,\mu$ m.

comes from the diffraction limit of the imaging optics, which is due to the numerical aperture of the aspheric lens, and was tested by removing the intensifier and illuminating a trapped atom for 100 ms. In this case, the atom acts as a point source for the imaging system and the associated response on the CCD is well fitted by a Gaussian shape with a size  $\sigma_{\text{diff}} = 0.5(1) \,\mu\text{m}$ .

The thermal distribution contributes for 0.3  $\mu$ m due to the transverse size of the distribution  $\sigma_{\perp}(0)$ , and 0.1  $\mu$ m due to the effect of depth of focus associated with the longitudinal size of the distribution  $\sigma_{\parallel}(0)$ . We measured this effect of the depth of focus by imaging a single trapped atom for various positions of the trap along the optical axis of the imaging system. Figure 4 shows the rms size of a Gaussian fit to the data, although for large values of the defocus  $\delta_z$  they slightly deviate from a Gaussian. The results tend asymptotically to the expected rms value of a disc with uniform intensity

$$\sigma_{\rm defocus} = \frac{1}{2} \,\delta z \times \tan \alpha,\tag{3}$$

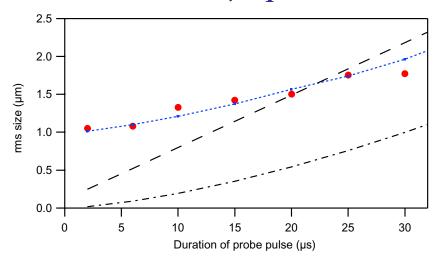
where  $\alpha$  is related to the numerical aperture by  $\sin \alpha = NA$ .

Finally, we analyze the contribution of the movement of the atom during the probe pulse. Firstly, the photon scattering by the probe induces a random walk of the atom, leading to an rms position in the plane perpendicular to the probe beam

$$\sigma_{\tau,\text{scatter}} = \frac{1}{3} v_{\text{rec}} \sqrt{R} \tau^{3/2}, \tag{4}$$

where  $v_{\rm rec}$  is the recoil velocity, *R* is the spontaneous emission rate and  $\tau$  is the duration of the probe pulse [15]. Secondly, the atom moves during the probe pulse due to the thermal velocity. An analytical calculation of the associated rms displacement yields

$$\sigma_{\tau,\text{thermal}} = \sigma_v \ \tau / \sqrt{3}. \tag{5}$$



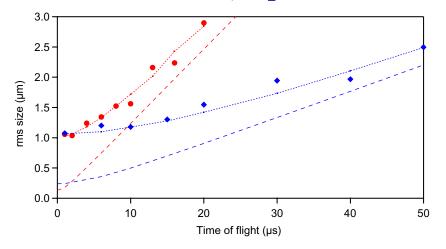
**Figure 5.** Effect on the measured spatial distribution of a single atom as a result of its movement during a probe pulse of variable length, after 1  $\mu$ s time-of-flight. The temperature of the atom is 150  $\mu$ K. Data (red circles) are in good agreement with the Monte Carlo simulation results (blue dotted line). At low pulse durations, the image size is limited by the intensifier response. Also shown are analytical rms contributions of equation (4) (dash dotted line) and of equation (5) (dashed line). The fact that the contribution of equation (5) is larger than both experimental and simulated results for long probe durations is due to the significant deviation of the distributions from a Gaussian model.

Both contributions (4) and (5) broaden the image of a single atom when the duration of the probe  $\tau$  is increased. We tested this effect by increasing  $\tau$  up to 30  $\mu$ s, as shown in figure 5. Although negligible for 2  $\mu$ s probe pulses and atoms at 150  $\mu$ K (as is the case in figure 5), this effect alone would be comparable to the intensifier response if we were using pulses as long as 20  $\mu$ s in the perspective of scattering more photons per shot and thus detect single atoms with a larger efficiency.

We confirmed the analysis above by a Monte Carlo simulation that takes into account all the effects mentioned above. It reproduces accurately our experimental data (see figures 4 and 5). Here, we note that the simulation indicates a significant deviation from a Gaussian shape for long probe durations or large values of the defocus. Because of the presence of noise in our imaging system, we did not consistently observe significant deviations on the real images and could not calculate any reliable value for the rms size of the images. We thus fitted our images with a Gaussian model and compared it to a Gaussian fit of our simulation. The discrepancy between the analytical expression (5) and the results in figure 5 is an indication of the error that we make by doing so. Note also that we have not included in the model the potential effect of the cooling of the atom by the counter-propagating probe beams.

#### 6. Temperature results and comparison to release-and-recapture experiments

We now come back to the temperature result obtained by fitting the data using equation (1). Among all the effects degrading the resolution, the depth of focus is the only one that



**Figure 6.** Time-of-flight measurements of a single atom for two different temperatures:  $T = 20(2) \,\mu\text{K}$  (blue diamonds) and  $T = 150(16) \,\mu\text{K}$  (red circles). The simulated results are also shown (blue and red dotted lines, respectively). The dashed lines (respectively blue and red) show the results of a simple model neglecting all broadening effects except for the transverse dimension of the thermal distribution,  $\sigma_{\perp}(0)$ .

varies with the time-of-flight as the atom can fly in the direction parallel to the optical axis. Therefore, equation (1) is not strictly valid in our case. We now use the Monte Carlo simulation mentioned above to fit the data shown in figure 3. The starting point of this simulation is a thermal distribution with a temperature that we adjust in order to reproduce the data. We find  $T = 20(2) \mu K$ . Not surprisingly, this result is in good agreement with the rough analysis mentioned in section 4, because the effect of the depth of focus is small.

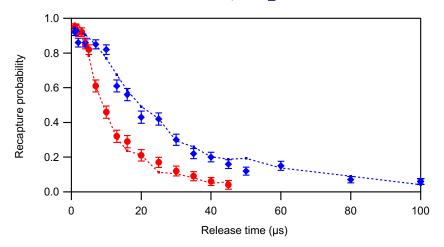
To cross check the temperature measurement, we use an independent method based on a release-and-recapture technique described in detail in [14]. This method uses an APD to detect the presence or the absence of the atom in the trap after release for a variable time. Figure 7 shows the results obtained by the release-and-recapture method. A fit to the data yields a temperature  $T = 19(2) \mu K$  in good agreement with the results of the time-of-flight method.

The two methods presented above can be extended to higher temperatures. For example, figures 6 and 7 show results obtained for a temperature  $T \sim 150 \,\mu\text{K}$ , achieved by leaving the dipole trap depth unchanged after loading it with a single atom. Our detection method is therefore applicable over a large range of temperatures with no anticipated limitation in the low-temperature range.

#### 7. Analysis of the noise of the imaging system

We now address the issue of the noise of our imaging system. The peak signal in the time-offlight image shown in figure 3(i) is 30 600 adu in 1 pixel<sup>6</sup> and corresponds to the detection of  $\sim$ 150 single atoms after 3400 shots of 2  $\mu$ s probe pulses. Normalized to one shot, the mean peak signal is thus 9 adu in 1 pixel. This should be compared to the background noise, which results

<sup>6</sup> 1 adu (analogue-to-digital unit) corresponds to the digitization step of the analogue signal acquired by the CCD camera. For the measurements presented in this paper, the electron-to-adu conversion factor was  $1 e^{-1} adu^{-1}$ .



**Figure 7.** Release-and-recapture measurements of a single atom for the same loading and cooling parameters as in figure 6. The fit (dotted lines) yields T = 19(2) and  $149(15) \mu K$  (data points are diamonds and circles, respectively).

from three contributions shown in figure 8: read-out noise from the CCD camera<sup>7</sup>, a background noise contribution from the probe light and a background noise contribution from spurious light (other than probe light).

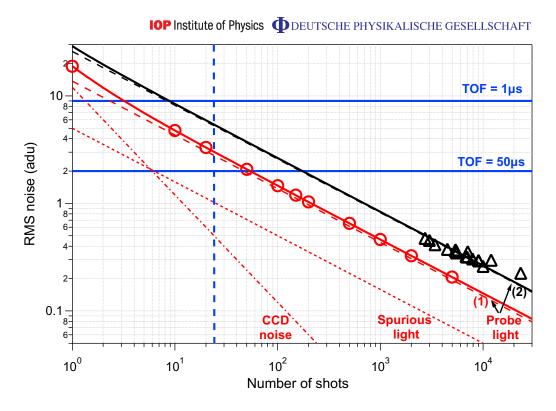
We have measured the read-out noise of the CCD camera and found 12 adu in one image. This noise is independent of the number of shots  $N_{seq}$  performed to acquire the image, because the CCD is read out only once after the probe pulses have illuminated the atom and the associated scattered light has fallen on the CCD. Normalized to one shot, the read-out noise of the CCD thus scales as  $12 \text{ adu}/N_{seq}$ . By contrast, the contributions, per shot, of the probe light and spurious light, scale as  $1/\sqrt{N_{seq}}$ . The three contributions were measured independently and add up quadratically to yield the data shown in figure 8. The signal-to-noise ratio is by far limited by the probe light contribution, which is due in part to scattering by atoms of the Rb beam intersecting the trapping region, and in part to scattering by the optics mounts under vacuum.

Figure 8 allows us to extract the number of sequences necessary to reach a given signalto-noise ratio. As explained at the end of section 4, the probability to detect one photon (and therefore one atom) in a single shot is  $4.4 \times 10^{-2}$  using a 2 µs duration probe, meaning that 23 shots are necessary to detect on average one photon. For a time-of-flight image to be correctly fitted, we have found that we need typically 100 detected photons, which implies 2300 sequences. For instance, in the case of the image shown in figure 3(i), the signal-to-noise ratio is ~20, whereas it is ~9 for figure 3(ii).

#### 8. Conclusion

In conclusion, we have demonstrated fluorescence imaging of a single atom in free flight by accumulating many images containing a single-photon event corresponding to a single atom.

<sup>7</sup> The dark count of the CCD being  $\sim 0.001$  adu pixel<sup>-1</sup> s<sup>-1</sup> is negligible for the parameters of our experiment.



**Figure 8.** Noise contributions, normalized to one shot, of our imaging system, versus the number of shots performed to acquire an image. The total rms noise (solid lines) is the quadratic sum of three contributions that were measured independently: the CCD read-out noise (red dash dotted line) and the background contributions due to spurious light (red dotted line) and probe light (dashed lines). The latter is larger when the atomic beam is switched on (case (2): black dashed line) than when it is off (case (1): red dashed line), due to the scattering by atoms of the atomic beam intersecting the probe beam. In both cases, the quadratic sum of the three contributions is in good agreement with the measured rms values of the total noise (black triangles and red circles, respectively). The peak signal, normalized to one shot (blue horizontal lines), is shown for two values of the time-of-flight,  $t_{\text{TOF}} = 1$  and 50  $\mu$ s, corresponding to images shown in figures 3(i) and (iii). The vertical dashed line indicates the minimum number of shots required to detect one atom.

We used this time-of-flight technique to measure the temperature of the atom after release from a microscopic optical dipole trap. This temperature measurement was confirmed by an independent method based on a release-and-recapture technique. The large numerical aperture of our imaging system and the extreme confinement of the atoms in the trap allow a high spatial resolution on the order of  $\sim 1 \,\mu$ m. The low noise level of our imaging system yields images showing  $\sim 150$  atoms with a very good signal to noise ratio ( $\sim 20$ ). These measurements have been performed in conditions where the atomic motion during the probe pulse can be completely neglected (see table 1). We thus obtained a very accurate characterization of the optical performance of our system. In a next step, it will be possible to increase the single-atom detection efficiency, by increasing the probe pulse duration (at the expense of a somehow reduced spatial resolution and signal-to-noise ratio). Finally, the measurements and the analysis presented in this work provide a calibration of our imaging system for future timeof-flight experiments where many atoms are confined in a microscopic dipole trap, and where interactions may play a central role.

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